

LECTURE 10

(10.1)

Acceleration as Fermi-Walker transport

- I. Moving vs. fixed frames
- II. Fermi-Walker transport
- III. Fermi - Walker transport of a vector

MTW [Ch. 6]

(10.2)

I. Moving Frame vs. Fixed Frame

The mathematization of accelerative change in the motion of particles/bodies gives rise to parametrized families of inertial reference frames, each one with an observer-defined tetrad of 4-vectors. They form a basis for the 4-d vector space whose origin is on the observer's world line. For an observer subject to uniform linear acceleration this world line is

$$X(\tau) = \{x^0(\tau), x^1(\tau), x^2(\tau), x^3(\tau)\} = \left\{ \frac{1}{g} \sinh \tau, \frac{1}{g} \cosh \tau, 0, 0 \right\},$$

and the corresponding parametrized family of Lorentz orthonormal frames is

$$u: (u)^{\mu} = (e_0)^{\mu} = (\cosh \tau, \sinh \tau, 0, 0)$$

$$\frac{du}{d\tau}: \left(\frac{du}{d\tau} \right)^{\mu} = (e_1)^{\mu} = (\sinh \tau, \cosh \tau, 0, 0)$$

$$e_2: (e_2)^{\mu} = (0, 0, 1, 0)$$

$$e_3: (e_3)^{\mu} = (0, 0, 0, 1)$$

$$\begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}$$

The components of this moving frame, $\{e_0(\tau), e_1(\tau), e_2(\tau), e_3(\tau)\}$, are typically given relative to a physically fixed basis

$$e_0: (1, 0, 0, 0)$$

$$e_1: (0, 1, 0, 0)$$

$$e_2: (0, 0, 1, 0)$$

$$e_3: (0, 0, 0, 1).$$

Its components are fixed because these basis vectors are aligned with the "fixed" stars.

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II. Fermi-Walker Transport

The physically essential property of the moving frame, Eq.(10.1), associated with a linear uniformly accelerated observer is that it is non-rotating. Three freely gimbaled gyroscopes aligned along the three spatial coordinate axes $\bar{x}, \bar{y}, \bar{z}$ in the rocket frame would not precess in the frame of the accelerating rocket as it executed its spacetime trajectory.

Fermi-Walker transport extends this non-rotation feature to a rocket subject to arbitrary acceleration.

I.) Fermi-Walker transport: Its construction.

A. GIVEN:

(i) The spacetime trajectory

$$X(\tau) : \{x^0(\tau), x^1(\tau), x^2(\tau), x^3(\tau)\}$$

of a particle/body subjected to some given acceleration.

(ii) The globally defined inertial LAB frame whose spatial axes e_1, e_2, e_3 point into the direction of some three chosen fixed stars

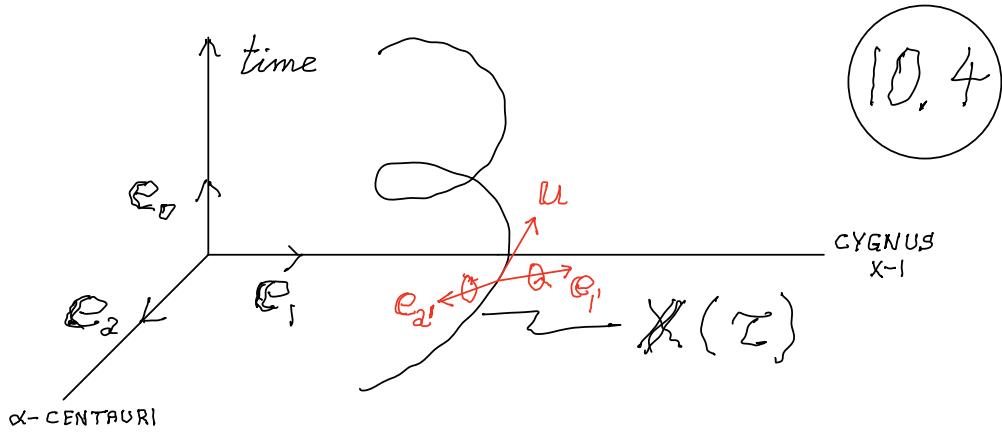


Figure 10.1: Inertial LAB frame with fixed orientation relative to the FIXED STARS vs an instantaneous Lorentz frame at a point event on $X(\tau)$.

(iii) Three GYROSCOPES attached to and carried along $X(\tau)$, the accelerated worldline, but gimbaled freely so that no torques act on each of the gyroscopes.

(iv) The τ -parametrized gyroscope-induced Lorentz orthonormal family of frames

$$\{E_{\mu}(\tau) = \frac{dX(\tau)}{d\tau} = u(\tau), E_1(\tau), E_2(\tau), E_3(\tau)\}$$

which consist of

a) the particle's 4-velocity

$$u(\tau) = \frac{dX}{d\tau} : \left\{ u^{\mu}(\tau) = \frac{dx^{\mu}}{d\tau}, \mu = 0, 1, 2, 3 \right\}$$

and

b) the three orthonormal directions determined by the three respective gyroscope directions.

10.5

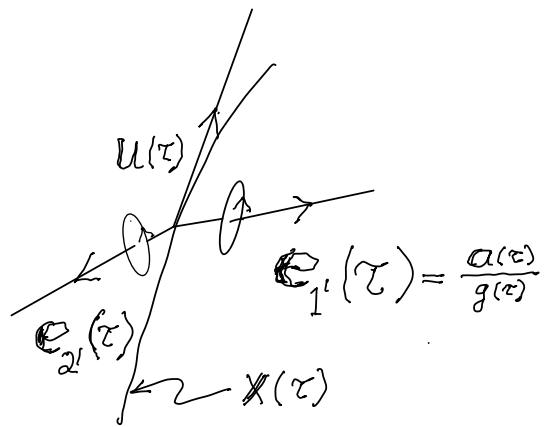


Figure 10.2: Gyroscope and 4-velocity determined Lorentz frame with one of the gyroscope axes aligned along the direction of the local acceleration vector α

(v) Notation:

Relative to the global LAB frame the gyroscope-determined 4-vectors have the components

$$\mathcal{E}_1 : \left\{ e_1^\mu : \mu = 0, 1, 2, 3 \right\}$$

$$\mathcal{E}_2 : \left\{ e_2^\mu : \mu = 0, 1, 2, 3 \right\}$$

$$\mathcal{E}_3 : \left\{ e_3^\mu : \mu = 0, 1, 2, 3 \right\}$$

$$u = \mathcal{E}_0 : \left\{ u^\mu : \mu = 0, 1, 2, 3 \right\}$$

vs.

relative to the comoving frame these same 4-vectors have the components

$$\mathcal{E}_1 : \left\{ e_1^\mu : \mu = 0, 1, 2, 3 \right\} = \{ 0, 1, 0, 0 \}$$

$$\mathcal{E}_2 : \left\{ e_2^\mu : \mu = 0, 1, 2, 3 \right\} = \{ 0, 0, 1, 0 \}$$

$$\mathcal{E}_3 : \left\{ e_3^\mu : \mu = 0, 1, 2, 3 \right\} = \{ 0, 0, 0, 1 \}$$

$$u = \mathcal{E}_0 : \left\{ u^\mu : \mu = 0, 1, 2, 3 \right\} = \{ 1, 0, 0, 0 \}$$

B. CONCLUSION:

(10.6)

- (i) The gyroscopes form a spatial platform with an orthonormal basis:

$$e_i \cdot e_j = e_i^{\mu} e_j^{\nu} \eta_{\mu\nu} = \eta_{ij} \left(= \begin{bmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix}; i, j = 1, 2, 3 \right)$$

- (ii) The time-like basis vector $e_0 = u$ is Lorentz normalized,

$$e_0 \cdot e_0 = u \cdot u = u^\mu u^\nu \eta_{\mu\nu} = -1$$

and Lorentz orthogonal to the spatial platform:

$$e_0 \cdot e_i = u \cdot e_i = 0 \quad i = 1, 2, 3$$

- (iii) Consequently,

$$e_\mu \cdot e_\nu = \eta_{\mu\nu} \left(= \begin{bmatrix} -1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix} \right)$$

is a set of 16 frame invariants, and

$$\{e_0(\tau), e_1(\tau), e_2(\tau), e_3(\tau)\} \equiv B_{FW}$$

is a τ -parametrized family of bases for the τ -parametrized family of instantaneous Lorentz frames (a.k.a. "moving frames")

- (iv) Summary:

The τ -parametrized family B_{FW} is determined by the freely gimbaled gyroscopes which are carried along a given accelerated worldline whose tangent is its instantaneous 4-velocity.

B_{FW} is called a Fermi-Walker basis

(10.7)

III. Fermi-Walker transport of a vector

Let $v(\tau)$ be a τ -parametrized family of vectors along a given curve $\chi(\tau)$ in spacetime

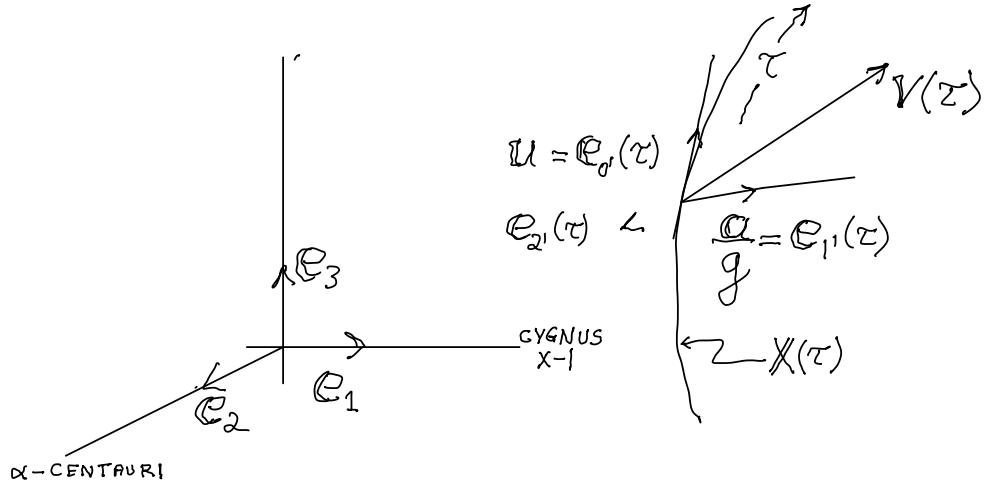


Figure 10.3: Instantaneous Fermi-Walker frame
in a global Lab (= "fixed stars") frame.

Expand $v(\tau)$ in terms of the moving Fermi-Walker basis and in terms of the fixed Lab basis. These expansions lead to the following

Definition (F-W Transport)

A τ -parametrized family of vectors $v(\tau)$ is said to be Fermi-Walker transported along the given spacetime trajectory $\chi(\tau)$ if its components remain constant relative to the gyroscope-induced basis

$$v(\tau) = \underbrace{e_{\mu}(\tau)}_{\text{constant}} \underbrace{v^{\mu}}_{\text{fixed (relative to the fixed stars)}} = \underbrace{e_{\mu}}_{\text{fixed (relative to the fixed stars)}} v^{\mu}(\tau)$$

Thus, a Fermi-Walker transported family of vectors $v(\tau)$ have components v^μ which are constant relative to the F-W basis $\{e_\mu(\tau)\}$,

$$\{e_\mu(\tau)\} : v(\tau) = e_\mu(\tau) v^\mu$$

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This gyroscope-induced basis changes relative to the fixed stars. On the other hand, $v(\tau)$ has components $v^\mu(\tau)$ which are changing relative to the fixed star basis $\{e_\mu\}$,

$$\{e_\mu\} : v(\tau) = e_\mu v^\mu(\tau)$$

IV. The law governing Fermi-Walker transport

The mathematization of the gyroscope-induced F-W transport consists of setting up the system of differential equations whose solutions are $v^\mu(\tau)$, the coefficients of $v(\tau)$ relative to the fixed star basis:

$$e_\mu v^\mu(\tau) = v(\tau) = e_\mu(\tau) v^\mu$$

MTW in their Section 6.5, p 170-171, show that this system of differential equations is linear and is given by

$$\frac{d v^\mu}{d \tau} = (u^\mu a^\nu - a^\mu u^\nu) v_\nu, \quad \mu = 0, 1, 2, 3$$

or in terms of frame invariant coefficients

$$\frac{d v^\mu}{d \tau} = u^\mu a \cdot v - a^\mu u \cdot v.$$

Here

$$u^\mu(\tau) = \frac{d x^\mu}{d \tau} \quad ("4\text{-velocity}")$$

$$a^\nu(\tau) = \frac{d^2 x^\nu}{d \tau^2} \quad ("4\text{-acceleration"})$$

$$v_\nu(\tau) = \eta_{\nu\sigma} v^\sigma(\tau)$$

Nota bene:

In MTW Ex. 6.9 is the problem of solving the F-W system for the case where $v(\tau) = S(\tau)$ is the spin 4-vector of a spinning electron in a circular orbit.

1. Notice that $v^\mu = u^\mu$ is a solution to the above system of equations.
2. Also notice that these equations imply that (i) with $v^\mu = u^\mu$ its rate of change $\frac{dv^\mu}{d\tau} = a^\mu$, and (ii) with $v^\mu = a^\mu$ its rate of change $\frac{dv^\mu}{d\tau} = g^2 u^\mu$.

This means that whatever changes u^μ and a^μ are subjected to, that these instantaneous changes are confined to the instantaneous plane spanned by u and a . There is no rotation out of this plane, i.e. the infinitesimal change in u^μ and a^μ is a "pure boost" confined to the instantaneous $u-a$ plane.