Lecture 23-24

Rotation On A 3-cube Immersed in A Curvature Field

Suggested reading: Chap. 2.8
In "Gravitation and Inertia"
by
J. Ciufolini & J.A. Wheeler

Consider a 3-dimensional domain D in 4-dimensional spacetime: $x^{\alpha}(u,v,t)$; $\alpha=0,1,2,3$; $\alpha\leq u\leq b$; $c\leq v\leq d$; $e\leq t\leq f$.

At each point-event P = {x*(u,v,t)} &D there are three vectors

tangent to D:

$$U = \frac{\partial x^{\alpha}}{\partial u} \frac{\partial}{\partial x^{\alpha}} = u^{\alpha} e_{\alpha}$$

$$V = \frac{\partial x^{\beta}}{\partial v} \frac{\partial}{\partial x^{\beta}} = v^{\beta} e_{\beta}$$

$$t = \frac{\partial x^{\delta}}{\partial t} \frac{\partial}{\partial x^{\delta}} = t^{\delta} e_{\delta}$$

Also at each such point consider a 3-d infinitesimal element of volume, a 3-cabe spanned by

$$\Delta U = \Delta U \frac{\partial x^{\alpha}}{\partial u} \frac{\partial}{\partial x^{\alpha}} = \Delta U U^{\alpha} e_{\alpha}$$

$$\Delta v V = \Delta v \frac{\partial x^{\beta}}{\partial v} \frac{\partial}{\partial x^{\beta}} = \Delta v v^{\beta} e_{\beta}$$

$$\Delta t t = \Delta t \frac{\partial x^{\gamma}}{\partial t} \frac{\partial}{\partial x^{\gamma}} = \Delta t t^{\gamma} e_{\gamma}$$

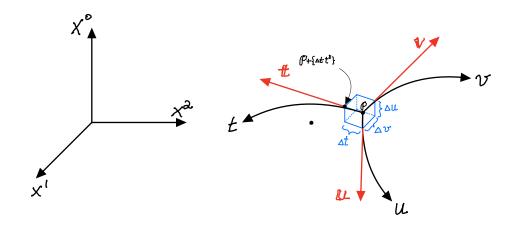


Figure 23.1 3-d cube in spacetime at point-event $P = \{x^{\alpha}(u,v,t)\}$ where it is spanned by coordinate increments $\Delta u, \Delta v,$ and Δt .

When such a 3-cube is permeated by a curvature field, each of its six faces features a <u>rotation</u>.

It vector W, upon being parallel transported around the boundary of each one of these faces, will be subjected to a rotational Change. Each change is proportional to the respective area of each face, to the strength of curvature field, and to the size and direction of that vector.

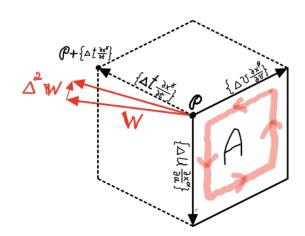


Figure 23.2 When a 3-cube is permeated by a curvature field, each of its faces features a rotation. It typical face in the figure is spanned by the two vectors, $\Delta U U = \Delta U \frac{\partial x}{\partial u} \frac{\partial x}{\partial x^{2}} \text{ and } \Delta v V = \Delta V \frac{\partial x}{\partial v} \frac{\partial x}{\partial x^{2}},$ both belonging to the vector space at P.

The rotation, which is featured by every face,

is mathematized by the bivector-valued 2-form

$$\vec{R}(,) = \underbrace{e_{\alpha \Lambda} e_{\beta}}_{2!} R^{\alpha \beta}_{\mu \nu} \underbrace{dx^{\mu}_{\lambda} dx^{\nu}}_{2!}(,).$$

The rotation $\mathcal{R}(\Delta u U, \Delta v V)$, when applied to the vector W, subjects it to the rotational change

$$\Delta^2 W = \mathcal{R}(\Delta u \mathcal{U}, \Delta v \mathcal{V}) \cdot W$$

$$= e_{\alpha} \mathcal{R}^{\alpha} \sigma_{\mu\nu} \mathcal{U}^{\mu} v^{\nu} w^{\sigma} \Delta u \Delta v$$

Whichever face features a rotation, it is mathematized by the bivector-valued 2-form,

$$\vec{R} = \frac{e_{x} \wedge e_{\beta}}{2!} R^{\alpha \beta}_{\mu \nu} \frac{d_{x}^{m} d_{x}^{\nu}}{2!},$$

This is a tensor of rank (2), a multilinear map.

It resulted, we recall, from evaluating the line integral

around the boundary ∂A of a 2-d domain A spanned by the vectors $\Delta u \mathbf{u} = \Delta u \frac{\partial x^{\alpha}}{\partial u} \frac{\partial}{\partial x^{\alpha}}$ and $\Delta v \mathbf{v} = \Delta v \frac{\partial x^{\beta}}{\partial v} \frac{\partial}{\partial x^{\beta}}$.

Use the infinitesimal 1-2 version of Stokes' theorem

$$\nabla_{\mathbf{u}} \vec{\Omega}(\mathbf{v}) - \nabla_{\mathbf{v}} \vec{\Omega}(\mathbf{u}) - \vec{\Omega}([\mathbf{u}, \mathbf{v}]) = d\vec{\Omega}(\mathbf{u}, \mathbf{v}), \qquad (23.1)$$

then Cartan's 2nd structure equation to calculate ds.

$$\Delta^2 W = \oint dW = d dW (\Delta u u, \Delta v v), \qquad (23.2)$$

and obtain
$$= e_{\alpha} R^{\alpha}_{\sigma \mu \nu} u^{\mu} v^{\nu} w^{\sigma} \Delta u \Delta v \qquad (23.3)$$

The r. h. s is the Riemann sum approximation consisting of only a single term for the double integral

$$\Delta^{2}W = \oint dW = \iint_{A} e_{\alpha} R^{\alpha}_{\sigma\mu\nu} u^{\mu} v^{\nu} w^{\sigma} du dv \qquad (23.4)$$

$$= \iint_{A} \frac{e_{\alpha} \wedge e_{\beta}}{2!} R^{\alpha\beta}_{\mu\nu} u^{\mu} v^{\nu} \cdot W du dv$$

$$= \iint_{A} R (U, V) \cdot W$$

over the infinitesimal 2-d domain A spanned by (\$u u, \$v v). With these two intermediates steps, Eqs. (23,2)-(23.3), understood, one writes Eq. (23.4) as

$$\Delta^2 W = \oint dw = \iint d(dw)$$

$$\partial A \qquad A$$

The infinitesimal version of the vectorial 1-2 Itohes' theorem holds for any combination of vectorial 1-forms. Without loss of generality let $\vec{\Omega} = \vec{f} dg$. This leads to the integral formulation of the 1-2 Stokes' theorem,

$$\int_{\partial A} \vec{\Omega} = \iint_{\mathbf{R}} d\vec{\Omega} \cdot \vec{\Omega}$$

II. Curvature-induced Rotational Change Due to the 6 Faces of a Cube

At 3-cube has 3 pairs of faces.

Q: What is the rotational change from all 3 pairs of opposing faces?

Ans: In 3 Steps:

Step1: From faces A' and A the rotational change is $\triangle^2 W_{A'} + \triangle^2 W_A =$

$$\oint_{\partial H} dw - \oint_{\partial H} dw = \begin{cases} \iint_{H^{1}} \vec{R} |_{\theta + \Delta t \cdot H} - \iint_{H} \vec{R} |_{\theta} \end{cases} \cdot w$$

$$= \begin{cases} \iint_{\theta + \Delta t \cdot H} \vec{R} (u, v) |_{\Delta t \cdot \Delta u \cdot \Delta v} |_{\theta} du dv \end{cases} \cdot w$$

$$= \begin{cases} \iint_{\theta + \Delta t \cdot H} \vec{R} (u, v) |_{\Delta t \cdot \Delta u \cdot \Delta v} \cdot w
\end{cases}$$

Step 2: Apply this mathematization to the other pairs of faces, B'&B and C'andC:

 $\Delta^2 \mathcal{W}_{A^1} + \Delta^2 \mathcal{W}_{A} + \Delta^2 \mathcal{W}_{B^1} + \Delta^2 \mathcal{W}_{B} + \Delta^2 \mathcal{W}_{C^1} + \Delta^2 \mathcal{W}_{C} =$

$$= \left[\int \int - \int + \int - \int + \int - \int \right] + \int - \int \\ \mathcal{R} \cdot w =$$

Each term is a covariant directional derivative of a sum of bivectors, tensors of rank $\binom{2}{6}$.

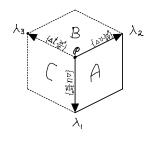
Step 3 Take advantage of the tensorial 2-3 version of Stokes' theorem

 $\nabla_{t} \vec{\Re}(u,v) + \nabla_{u} \vec{\Re}(v,t) + \nabla_{v} \vec{\Re}(t,u)$ $-\vec{\Re}([u,v],t) - \vec{\Re}([v,t],u) - \vec{\Re}([t,u],v) = d\vec{\Re}(u,v,t)$

and

Let $\mathcal{D} = \{ x^{\alpha}(\Delta u, \Delta v, \Delta t) \} = \{ x^{\alpha} = C(\lambda_1, \lambda_2, \lambda_3) : o \leq \lambda_1, \lambda_2, \lambda_3 \leq 1 \}$

be the interior of the 3-cube



Let 2D = oriented 2-d boundary of D, the 3-cube's oriented interior spanned by t, u, and v.

2D consists of the six oriented faces of D

$$C^{1} = C^{1+} = C(1, \lambda_{2}, \lambda_{3})$$
 right(back)
 $C = C^{1-} = C(0, \lambda_{2}, \lambda_{3})$ left (front)
 $B^{1} = C^{2+} = C(\lambda_{1}, 1, \lambda_{3})$ upper
 $B = C^{2-} = C(\lambda_{1}, 0, \lambda_{3})$ lower
 $A^{1} = C^{3+} = C(\lambda_{1}, \lambda_{2}, 1)$ left (back)
 $A = C^{3-} = C(\lambda_{1}, \lambda_{2}, 0)$ right (front)

The oriented boundary 2D of D is

$$\partial D = \sum_{j=1}^{3} (-1)^{j-1} (c^{j+1} - c^{j-1})$$

$$= (c^{1+1} - c^{1-1}) - (c^{2+1} - c^{2-1}) + (c^{3+1} - c^{3-1})$$

$$= (c^{1} - c^{1}) - (B^{1} - B) + (A^{1} - A)$$

$$\partial D = \sum_{l=1}^{6} [l^{tR} Face] = (A^{1} - A) - (B^{1} - B) + (c^{1} - c)$$

Appendix

Cartan's 2^{nd} structure equation $ddw = d\vec{s}\vec{z} = e_{\omega}(d\omega^{\kappa}_{\rho} + \omega^{\kappa}_{\gamma} \wedge \omega^{\kappa}_{\rho})w^{\beta} = e_{\omega}R^{\kappa}_{\rho,n}\frac{d\kappa^{\kappa}_{\rho}d\kappa^{\kappa}}{2!}w^{\beta}$ is linked to the physical world through its evaluation

Evaluate it on two spanning vectors of a face, say
and and $\Delta v V$ depicted in Figure 23,2. Then apply it to
the vector $W = e_{\sigma} w^{\sigma}$,

The result is a rotational change which is enjoyed by W. This change is $\Delta^{2}W = \hat{R}(\Delta u u, \Delta v v) \cdot W = \frac{\vec{e}_{\omega} \wedge \vec{e}_{\beta}}{2!} R^{\alpha\beta} \frac{dx^{\kappa} \wedge dx^{\kappa}}{2!} (\Delta u u, \Delta v v) \cdot W.$

Before evaluating it, ask and answer the following question: Why is it rotational? The answer is this: by evaluating it, one sees that it is the sum of rotations in each of the $e_{\kappa}-e_{\beta}$ planes,

$$\frac{\hat{R}}{\hat{R}} (\Delta u \, u, \Delta v \, v) = \frac{e_{u} \wedge e_{\beta}}{2!} \, \hat{R}^{\alpha\beta} (\Delta u \, u, \Delta v \, v) \qquad (23.1)$$

$$= \frac{e_{u} \wedge e_{\beta}}{2!} \, \hat{R}^{\alpha\beta} (\Delta u \, u, \Delta v \, v) \qquad (23.1)$$

$$= \frac{e_{u} \wedge e_{\beta}}{2!} \, \hat{R}^{\alpha\beta} (\Delta u \, u, \Delta v \, v) \qquad (23.1)$$

The angular amount of this rotation, which is induced by the curvature, is

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This is mathematically the same as described in Lecture 19 on page 19.10, but with a physical difference. Indeed, there Eqs.(19.2) and (19.3) refer instead to rotations which are induced by the angular velocity $\vec{w} = e_{\mathbf{k}} \, \omega^{\mathbf{k}}$ with

O'j = E'j & what

as the angle of rotation in the e_i - e_j plane

The evaluation of the curvature-induced change
in the vector w is achieved by recalling that $dx^{\mu}(u) = u^{\mu}$ and that $e_{\beta} \cdot w = e_{\beta} \cdot e_{\sigma} w = g_{\beta\sigma} \cdot w^{\sigma}$. The result is $dx^{\mu}(u) = e_{\alpha} R^{\alpha}_{\sigma\mu\nu} u^{\mu} v^{\nu} w^{\sigma} \Delta u \Delta v$.

But this is only the contribution to the vectorial change in $W = e_{\sigma} w^{\sigma}$ which comes from that face of the 3-cube which is spanned by $\Delta u u$ and $\Delta v v$ in Figure 23,2.

$$F = F_{\mu\nu} dx^{\mu} dx^{\nu}/2!$$

permeating spacetime and (ii) a two-dimensional

surface A parametrized by u and v, $x^{\alpha}(u, v)$.

The vectors tangent to A are

$$U = \frac{\partial x^{\alpha}}{\partial u} \frac{\partial}{\partial x^{\alpha}} = u^{\alpha} e_{\alpha}, \quad V = \frac{\partial x^{\beta}}{\partial v} \frac{\partial}{\partial x^{\beta}} = v^{\beta} e_{\beta}$$

The surface integral of E over A is mathematized by

$$\iint_{A} \mathbf{F} = \iint_{A} \mathbf{F}_{\mu\nu} dx^{\mu} dx^{\nu}/2! \qquad (23.3)$$

 $\equiv \iint_{A} F(u, v) du dv$

 $=\iint_{\Lambda} F_{\mu\nu} \left(x^{\sigma}(u,\nu) \right) dx^{\mu} dx^{\nu}/2! \left(U,V \right) du d\nu$

 $=\iint\limits_{A}F_{\mu\nu}(x^{s}(u,\nu))\ u^{\mu}v^{\nu}du\ dv$

When the integration domain surrounding a point $P \in A$ is so small that $F_{\mu\nu}$ is constant to lowest order, the value of the integral in the neighborhood of P is

$$\iint_{A} \mathbf{F} = \iint_{A} \mathbf{F}_{\mu\nu} (\mathbf{x}^{\sigma}(\mathbf{u}, \mathbf{v})) \quad \mathbf{u}^{\mu} \mathbf{v}^{\nu} \Delta \mathbf{u} \Delta \mathbf{v}$$

$$\iint_{A} F = \iint_{A} F_{\mu\nu} dx^{\mu} dx^{\nu} dx^{\nu$$